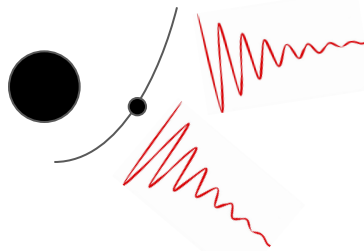


Motivation: EMRI waveforms

- LISA is expected to detect EMRI signals modeled by Black hole perturbation theory.



- The Teukolsky master equation describes scalar, vector, and tensor field perturbations in the space-time of Kerr black holes and given by

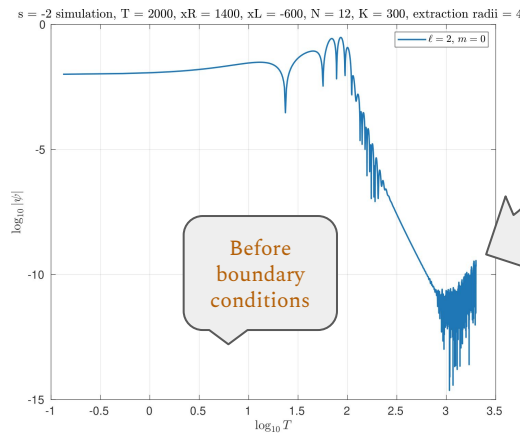
$$\begin{aligned}
 & - \left[\frac{(r^2 + a^2)^2}{\Delta} - a^2 \sin^2 \theta \right] \partial_{tt} \Psi - \frac{4Mar}{\Delta} \partial_{t\phi} \Psi \\
 & - 2s \left[r - \frac{M(r^2 - a^2)}{\Delta} + ia \cos \theta \right] \partial_t \Psi \\
 & + \Delta^{-s} \partial_r (\Delta^{s+1} \partial_r \Psi) + \frac{1}{\sin \theta} \partial_\theta (\sin \theta \partial_\theta \Psi) + \\
 & \left[\frac{1}{\sin^2 \theta} - \frac{a^2}{\Delta} \right] \partial_{\phi\phi} \Psi + 2s \left[\frac{a(r-M)}{\Delta} + \frac{i \cos \theta}{\sin^2 \theta} \right] \partial_\phi \Psi \\
 & - (s^2 \cot^2 \theta - s) \Psi = -4\pi (r^2 + a^2 \cos^2 \theta) T,
 \end{aligned}$$

$$\Delta = r^2 - 2Mr + a^2$$

- s is the spin weight of the field, where the $s = \pm 2$ versions of these equations describe gravitational waves, and T is a source term.
- We solve the Teukolsky equation using **discontinuous Galerkin** methods.

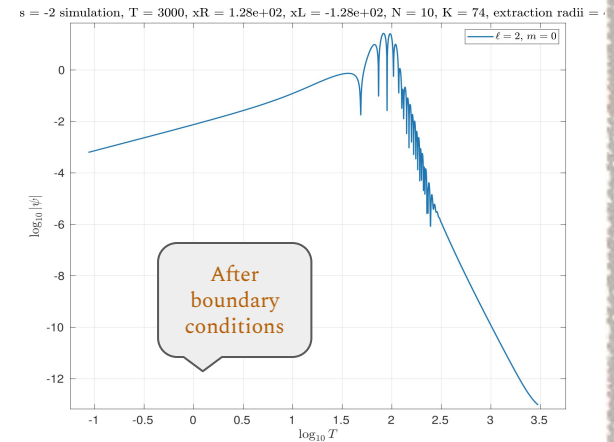
Problems in Boyer-Lindquist coordinates ?

Spurious growth at $\sim (100M)$

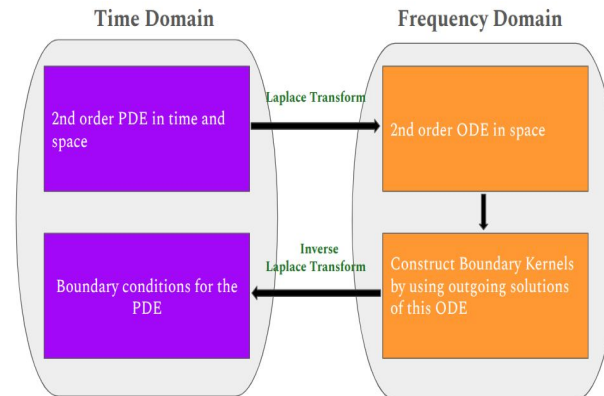


Exact BCs resolve Blow up

Simulation with exact boundary conditions applied: ✓



Outer Boundary conditions using Boundary Kernels



Validation using power-law decay of tails: ✓

